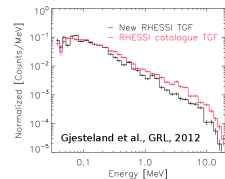
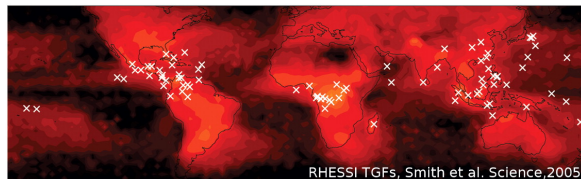


# Introduction: Runaways electrons and TGFs

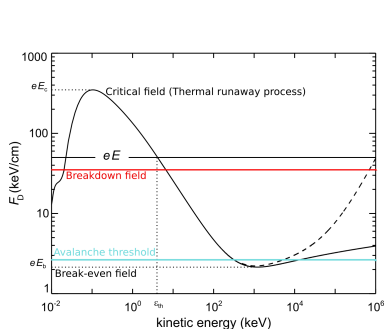
- C.T.R. Wilson 1925: Runaway electrons (**RE**), bremsstrahlung.
- Terrestrial gamma-ray flashes (**TGFs**): *Fishman et al., Science 264 (1994)*.  
BATSE: 76 detections of **TGFs** in 9 years.  
RHESSI: >500 detections (2004–2006).
- Possibly >50 **TGFs** per day (>3M lightning events).
- Source altitude < 20 km, energy of photons up to 20 MeV.
- **Accepted scenario**: bremsstrahlung of runaway electrons.



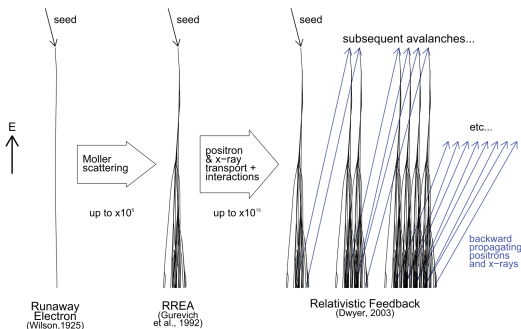
**Detailed mechanism by which lightning is associated with TGFs is still a mystery.**

# Runaway electrons

- Dynamic friction force (Rate of energy loss):  $\frac{\partial p}{\partial t} = eE - F$ .
- Runaway electrons in air: Wilson 1925.
- Relativistic Runaway Electron Avalanche (**RREA**): Gurevich 1992;
- Relativistic feedback: Dwyer 2003 (Avalanche threshold  $E_{th} = 2.84 \times 10^5$  V/m)

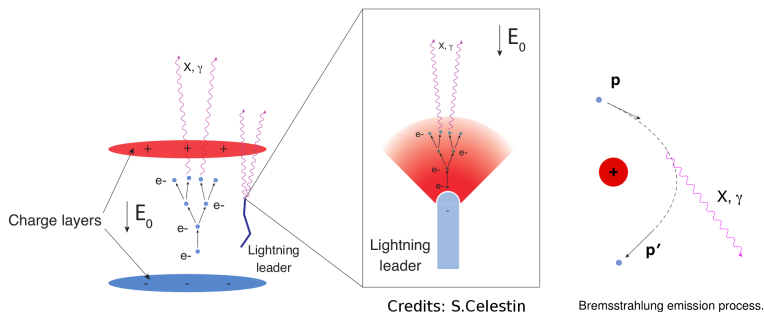


(Fig. source: Dwyer et al., *Space Sci Rev*, 2012.)



## Two theories to explain TGFs

- Relativistic Runaway Electron Avalanche + Feedback (e.g., Dwyer)  
→ high energy seed electron needed.
- Thermal runaway process (e.g., Moss 2006, Chanrion 2008, 2014, Li 2012, Celestin 2011).  
→ **acceleration of electrons in streamer fields** (tens of kV).  
→ further acceleration in the field of leaders (tens of MV).



**Needed:**  $10^{17}$  runaway electrons,  $>1\text{MeV}$ , (Dwyer, GRL, 2003).

**Is the thermal runaway process from streamers able to provide enough of RE??**

# Introduction: What is a **streamer**?

## Streamer breakdown theory (Loeb and Meek 1930')

- 'Fast' breakdown for spark discharge at atm pressure
- Time scale  $< 10^{-7}$  s for one-cm gap:
  - ions too slow to be blamed;
  - (classical theory relied on secondary emission)
- New concept: 'Streamers'

## Streamers:

Contracted ionizing waves with self-generated field enhancement that propagates into a low-ionized medium exposed to high electric field leaving filamentary trails of plasma behind.

Characteristics:

**Velocity** even 1% speed of light;

**Radius** typically  $10^{-4}$ – $10^{-3}$  m at STP;

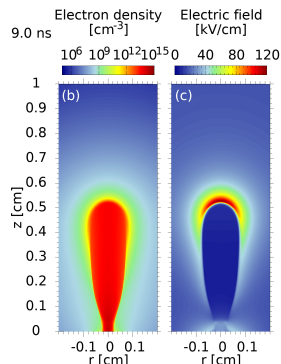
**Peak electric field** at Streamer head SH (net space charge):  $4$ – $7 E_k$  ( $E_k \sim 32$  kV/cm at STP)

↪ **Intense electron impact ionization**: space charge region move forward;

**Polarity** (sign of charge in the streamer head):

- **(+) Positive**: propagate. against electron drift: need of ambient seed charges
- **(-) Negative**: el. aval. from SH ↪ propagate w/o need of seed electrons

[R. V. Hodges, 1985 Phys. Rev. A; Raizer, 1991; V. Pasko 2006, U. Ebert 2006 PSSST]



# Introduction: Why are streamers interesting?

C.T.R. Wilson (1925): Possibility of large scale discharges high above thunderclouds.

if(

The dipole electric field of the thunderclouds falls with altitude as  $1/h^3$ .

.and.

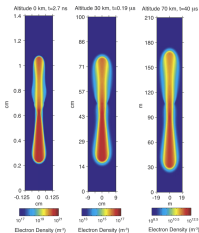
Critical breakdown field falls exponentially with decreasing air density.

)then

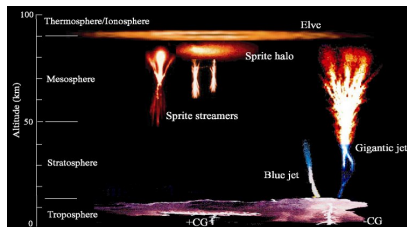
There must be an altitude high above where the electric force would exceed the sparking limit.

endif

Double headed streamers  $\rightsquigarrow$  precursors to more complicated phenomena.



[V.P. Pasko PSST 2007]



[credits: cnes]

Since 1989 many Transient Luminous Events discovered.

**But streamers play a role even for more unusual phenomena!**



# Modeling of streamers: Classical fluid model

How to model the streamer discharge?

It depends on the physical question we want to address!

**Basic ingredients** [U. Ebert, PSST, 2006; A. Luque, U. Ebert, JCP 2012]:

- (1) Ionization processes (electron impact, photons)
- (2) Drift and diffusion of charged particles in electric field
- (3) Effects of a space charge on the electric field

**'Classical' fluid streamer model:**

$$\begin{aligned} \partial_t n_s - \text{sign}(q_s) \nabla \cdot (n_s \mathbf{v}_s) - \nabla \cdot (D_s \nabla n_s) &= S_s, \\ \varepsilon_0 \nabla^2 \varphi &= \sum_s q_s n_s, \quad s \in \{\text{charged species}\}. \end{aligned}$$

**Approximations:**

- (a) Density Approximation: distribution of charged species described with densities
- (b) Local Field Approximation: Electrons relax rapidly enough  
 $\rightsquigarrow$  in equilibrium with  $\mathbf{E}$  at given place.
- (c) Electrostatic Approximation: Induced electric fields are negligible ( $\mathbf{E} = -\nabla\varphi$ ).

**Challenges:** spatial and time scales span many orders of magnitude.

**Classical model is not sufficient to study electron acceleration!**

# Hybrid model

A beam-bulk approach proposed by Belenguer and Boeuf for RF discharge (Phys. Rev. A, 1990). Model combines 2 models for 2 electron groups divided in the energy space:

## Electrons below 100 eV

### Axisymmetric fluid streamer model in a point to plain geometry

[Pecherau2012,Celestin2009]

- Classical drift-diffusion model based on [Kulikovsky1997]
- The mobility, diffusion coefficient and reaction rates are calculated from Monte Carlo model for electrons of energy below 100 eV.
- Electric field is calculated solving the Poisson equation

## Electrons above 100 eV

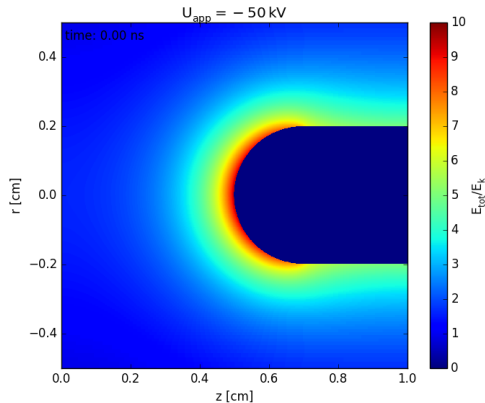
### 2D-s,3D-V PIC-MCC model

- Standard PIC-MCC code follows the electron trajectories in the electric field and their collisions with neutral air molecules [Chanrion, JCP, 2008] .
- Injection of electrons Fluid  $\rightarrow$  PIC based on fluid macroscopic parameters using precomputed distribution functions.
- Electrons that slow below 100 eV are reinjected in the fluid model.

- Coupling between the two models is repeated every time step.

# Test case

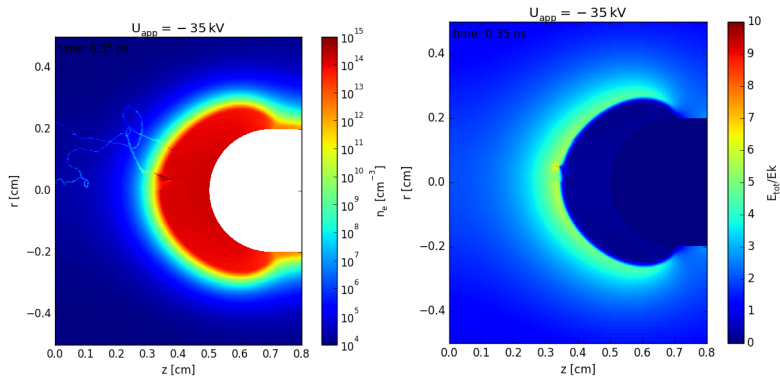
- Hemisphere pin electrode (0.2 cm radius).
- 0.5 cm gap.
- Applied voltage: (negative polarity) step function with amplitude:  $\{-35, -45, -50\text{ kV}\}$
- Air at atmospheric pressure
- Low preionization of  $10^4 \text{ cm}^{-3}$



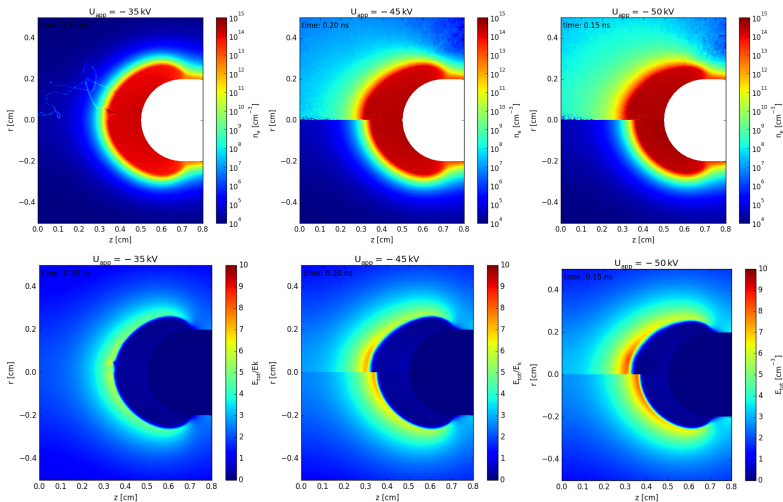
**Photoionization not included: to accent the role fast electrons**



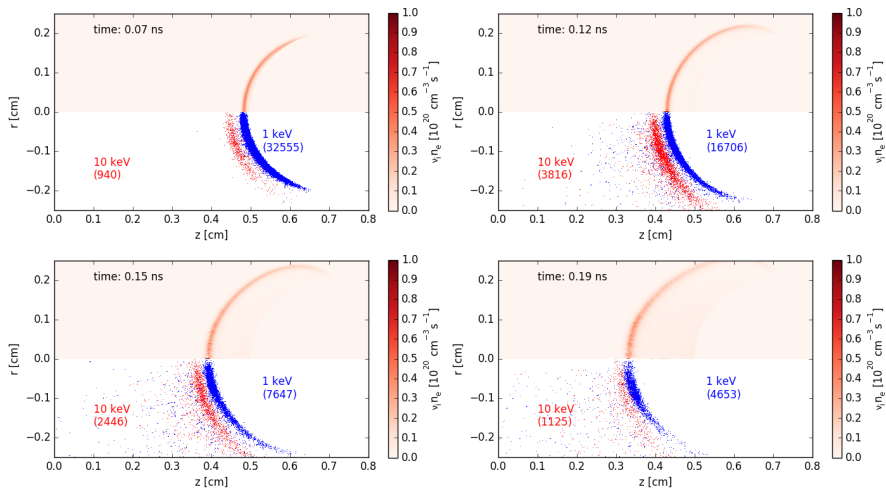
# Effect of fast electrons on discharge: -35kV



# Effect of fast electrons on discharge



## RE acceleration source



## Electron acceleration: z-energy space

